



3. Freezing Rotating Waves in Reaction Diffusion Systems

3.1 Rotating waves in reaction diffusion systems

Consider a [system of reaction diffusion equations](#) in d space dimensions

$$(1) \quad \begin{aligned} u_t(x, t) &= A\Delta u(x, t) + f(u(x, t)) \quad , \quad x \in \mathbb{R}^d, t > 0, d \geq 2, \\ u(x, 0) &= u_0(x) \quad , \quad x \in \mathbb{R}^d, t = 0, \end{aligned}$$

with [diffusion matrix](#) $A \in \mathbb{R}^{m,m}$, [nonlinearity](#) $f : \mathbb{R}^m \rightarrow \mathbb{R}^m$, [initial data](#) $u_0 : \mathbb{R}^d \rightarrow \mathbb{R}^m$ and [solution](#) $u : \mathbb{R}^d \times [0, \infty) \rightarrow \mathbb{R}^m$. The operator Δ denotes the Laplacian given by

$$\Delta u(x) := \sum_{i=1}^d \frac{\partial^2 u}{\partial x_i^2}(x), \quad x \in \mathbb{R}^d.$$

We are interested in rotating wave solutions of (1): A [rotating wave](#) of (1) is a solution $u_\star : \mathbb{R}^d \times [0, \infty) \rightarrow \mathbb{R}^m$ of the form

$$(2) \quad u_\star(x, t) = v_\star(e^{-tS_\star}(x - x_\star)) \quad , \quad x \in \mathbb{R}^d, t \geq 0,$$

with a time-independent function $v_\star : \mathbb{R}^d \rightarrow \mathbb{R}^m$, a skew-symmetric matrix $S_\star \in \mathbb{R}^{d,d}$, i.e. $S_\star^T = -S_\star$, and a vector $x_\star \in \mathbb{R}^d$. The function $v_\star : \mathbb{R}^d \rightarrow \mathbb{R}^m$ is called the [profile](#) of the rotating wave. The vector $x_\star \in \mathbb{R}^d$ can be considered as the [center of rotation](#) for $d = 2$ and as the [support vector](#) of the axis of rotation for $d = 3$. In case $d \in \{2, 3\}$, $S_\star \in \mathbb{R}^{d,d}$ can be considered as the [angular velocity tensor](#) associated to the angular velocity vector $\omega \in \mathbb{R}^{\frac{d(d-1)}{2}}$ containing S_\star^{ij} , $i = 1, \dots, d-1$, $j = i+1, \dots, d$. Note that the skew-symmetry of S_\star implies that e^{-tS_\star} is a rotation matrix. A rotating wave u_\star is called a [localized](#), if

$$\lim_{|x| \rightarrow \infty} |v_\star(x) - v_\infty| = 0 \text{ for some } v_\infty \in \mathbb{R}^m.$$

Our aim is to approximate rotating wave solutions of (1). The idea for approximating the rotating wave u_\star is to determine the profile v_\star , the angular velocity tensor S_\star and the vector x_\star . This requires to transform (1) into a co-rotating coordinate system.

Transforming (1) via $u(x, t) = v(\xi, t)$ with $\xi := e^{-tS_\star}(x - x_\star)$ in a [co-rotating frame](#) yields

$$(3) \quad \begin{aligned} v_t(\xi, t) &= A\Delta v(\xi, t) + \langle S_\star \xi, \nabla v(\xi, t) \rangle + f(v(\xi, t)) \quad , \quad \xi \in \mathbb{R}^d, t > 0, d \geq 2, \\ v(\xi, 0) &= u_0(\xi) \quad , \quad \xi \in \mathbb{R}^d, t = 0, \end{aligned}$$

where the [drift term](#) is defined by

$$(4) \quad \langle S_\star x, \nabla v(x) \rangle := \sum_{i=1}^d \sum_{j=1}^d S_\star^{ij} x_j D_i v(x), \quad D_i := \frac{\partial}{\partial x_i}$$

In case of skew-symmetric matrices $S_\star = -S_\star^T$, the drift term from (3) is a [rotational term](#) containing angular derivatives

$$(5) \quad \langle S_\star x, \nabla v(x) \rangle = \sum_{i=1}^{d-1} \sum_{j=i+1}^d S_\star^{ij} (x_j D_i - x_i D_j) v(x).$$

The operator $A\Delta v(x) + \langle S_*x, \nabla v(x) \rangle$ is called the **Ornstein-Uhlenbeck operator**. It is given by the sum of the diffusion term $A\Delta v(x)$ and the drift term $\langle S_*x, \nabla v(x) \rangle$ and has unbounded (indeed linearly growing) coefficients.

Inserting (2) into (1) shows, that v_* is a stationary solution of (3), i.e. v_* solves the **rotating wave equation**

$$(6) \quad 0 = A\Delta v_*(\xi) + \langle S_*\xi, \nabla v_*(\xi) \rangle + f(v_*(\xi)) \quad , \xi \in \mathbb{R}^d.$$

We are also interested in nonlinear stability of rotating waves. It is well known from the literature, that (at least for $d = 2$) spectral stability implies nonlinear stability. For investigating spectral stability of rotating waves, we must analyze the spectrum of the linearization of the right hand side in (3) at the wave profile v_* , i.e.

$$[\mathcal{L}w](\xi) = A\Delta w(\xi) + \langle S_*\xi, \nabla w(\xi) \rangle + Df(v_*(\xi))w(\xi) \quad , \xi \in \mathbb{R}^d.$$

This requires to find solutions (λ, w) with $\lambda \in \mathbb{C}$ and $w : \mathbb{R}^d \rightarrow \mathbb{C}^m$ of the **eigenvalue problem**

$$(7) \quad \lambda w(\xi) = A\Delta w(\xi) + \langle S_*\xi, \nabla w(\xi) \rangle + Df(v_*(\xi))w(\xi) \quad , \xi \in \mathbb{R}^d,$$

with **eigenfunction** $w : \mathbb{R}^d \rightarrow \mathbb{C}^m$ and **eigenvalue** $\lambda \in \mathbb{C}$.

Approximating v_* via (3) requires the knowledge about the angular velocity tensor S_* which is in general unknown. This motivates to introduce the freezing method, whose idea is it to approximate the profile v_* , the angular velocity tensor S_* and the vector $x_* \in \mathbb{R}^d$ simultaneously.

3.2 Freezing method for rotating waves

Consider again a **system of reaction diffusion equations**, cf. (1),

$$(8) \quad \begin{aligned} u_t(x, t) &= A\Delta u(x, t) + f(u(x, t)) \quad , x \in \mathbb{R}^d, t > 0, d \geq 2, \\ u(x, 0) &= u_0(x) \quad , x \in \mathbb{R}, t = 0. \end{aligned}$$

Introducing new unknowns $\gamma(t) = (R(t), \tau(t)) \in \text{SO}(d) \times \mathbb{R}^d = \text{SE}(d)$ (**position**) and $v(\xi, t) \in \mathbb{R}^m$ (**profile**) via the **rotating wave ansatz**

$$(9) \quad u(x, t) = v(\xi, t), \xi := R(t)^{-1}(x - \tau(t)) \quad , x \in \mathbb{R}^d, t \geq 0,$$

inserting (9) into (8) (the computation will be omitted) and introducing a new unknown $\mu(t) = (S(t), \lambda(t)) \in \mathfrak{so}(d) \times \mathbb{R}^d = \mathfrak{se}(d)$ via $\begin{pmatrix} R_t(t) \\ \tau_t(t) \end{pmatrix} = \begin{pmatrix} R(t)S(t) \\ R(t)\lambda(t) \end{pmatrix}$ yields

$$(10) \quad \begin{aligned} v_t(\xi, t) &= A\Delta v(\xi, t) + \langle S(t)\xi + \lambda(t), \nabla v(\xi, t) \rangle + f(v(\xi, t)) \quad , \xi \in \mathbb{R}^d, t > 0, \\ \begin{pmatrix} R_t(t) \\ \tau_t(t) \end{pmatrix} &= \begin{pmatrix} R(t)S(t) \\ R(t)\lambda(t) \end{pmatrix} \quad , t > 0. \end{aligned}$$

Equ. (10) has to be equipped with suitable initial data. Requiring $R(0) = I_d$ and $\tau(0) = 0$, (9) and (8) imply

$$(11) \quad v(\xi, 0) = u_0(\xi) \quad , \xi \in \mathbb{R}^d, t = 0.$$

Collecting the equations (10), $R(0) = I_d$, $\tau(0) = 0$ and (11) we obtain

$$(12) \quad \begin{aligned} v_t(\xi, t) &= A\Delta v(\xi, t) + \langle S(t)\xi + \lambda(t), \nabla v(\xi, t) \rangle + f(v(\xi, t)) \quad , \xi \in \mathbb{R}^d, t > 0, \\ v(\xi, 0) &= u_0(\xi) \quad , \xi \in \mathbb{R}^d, t = 0, \\ \begin{pmatrix} R_t(t) \\ \tau_t(t) \end{pmatrix} &= \begin{pmatrix} R(t)S(t) \\ R(t)\lambda(t) \end{pmatrix} \quad , t > 0, \\ \begin{pmatrix} R(0) \\ \tau(0) \end{pmatrix} &= \begin{pmatrix} I_d \\ 0 \end{pmatrix} \quad , t = 0. \end{aligned}$$

(12) contains the equations for v and $\gamma = (R, \tau)$. But so far, the system (12) is not well-posed, since there is still no equation for $\mu = (S, \lambda)$. To determine μ we require an additional algebraic

constraint, a so called phase condition: For this purpose let $\hat{v} : \mathbb{R}^d \rightarrow \mathbb{R}^m$ be a template function, e.g. $\hat{v} = u_0$. The idea of the [phase condition](#) is to choose $v(\cdot, t)$ such that

$$\min_{g:=(\tilde{R}, \tilde{\tau}) \in \mathfrak{SE}(d)} \left\| v(\cdot, t) - \hat{v}(\tilde{R}^{-1}(\cdot - \tilde{\tau})) \right\|_{L^2(\mathbb{R}^d, \mathbb{R}^m)}^2 = \|v(\cdot, t) - \hat{v}(\cdot)\|_{L^2(\mathbb{R}^d, \mathbb{R}^m)}^2, \quad t \geq 0.$$

A necessary condition to guarantee that the left hand side attains its minimum at $g = (I_d, 0)$ is that the first derivative of $\left\| v(\cdot, t) - \hat{v}(\tilde{R}^{-1}(\cdot - \tilde{\tau})) \right\|_{L^2(\mathbb{R}^d, \mathbb{R}^m)}^2$ evaluated at $g := (\tilde{R}, \tilde{\tau}) = (I_d, 0)$ vanishes, i.e.

$$(13) \quad \begin{aligned} 0 &\stackrel{!}{=} \left[\frac{d}{d(\tilde{R}, \tilde{\tau})} \left(v(\cdot, t) - \hat{v}(\tilde{R}^{-1}(\cdot - \tilde{\tau})), v(\cdot, t) - \hat{v}(\tilde{R}^{-1}(\cdot - \tilde{\tau})) \right)_{L^2(\mathbb{R}^d, \mathbb{R}^m)} \right]_{g=0} \\ &= 2 \left(v(\cdot, t) - \hat{v}, \left\langle \tilde{S} \cdot + \tilde{\lambda}, \nabla \hat{v} \right\rangle \right)_{L^2(\mathbb{R}^d, \mathbb{R}^m)} \quad \forall (\tilde{S}, \tilde{\lambda}) \in \mathfrak{se}(d). \end{aligned}$$

Requiring (13) for every basis element $(I_{ij} - I_{ji}, 0)$ ($i = 1, \dots, d-1, j = i+1, \dots, d$) and $(0, e_l)$ ($l = 1, \dots, d$) of $\mathfrak{se}(d)$, we obtain for all $t \geq 0$

$$(14) \quad \begin{aligned} 0 &= (v(\cdot, t) - \hat{v}, (\xi_j D_i - \xi_i D_j) \hat{v})_{L^2(\mathbb{R}^d, \mathbb{R}^m)}, \quad i = 1, \dots, d-1, j = i+1, \dots, d, \\ 0 &= (v(\cdot, t) - \hat{v}, D_l \hat{v})_{L^2(\mathbb{R}^d, \mathbb{R}^m)}, \quad l = 1, \dots, d. \end{aligned}$$

Combining (12) and (14) yields a [partial differential algebraic evolution equation \(PDAE\)](#)

$$(15) \quad \begin{aligned} v_t(\xi, t) &= A \Delta v(\xi, t) + \langle S(t) \xi + \lambda(t), \nabla v(\xi, t) \rangle + f(v(\xi, t)) \quad , \quad \xi \in \mathbb{R}^d, t > 0, \\ v(\xi, 0) &= u_0(\xi) \quad , \quad \xi \in \mathbb{R}^d, t = 0, \\ 0 &= (v(\cdot, t) - \hat{v}, (\xi_j D_i - \xi_i D_j) \hat{v})_{L^2(\mathbb{R}^d, \mathbb{R}^m)} \quad , \quad t \geq 0, \\ 0 &= (v(\cdot, t) - \hat{v}, D_l \hat{v})_{L^2(\mathbb{R}^d, \mathbb{R}^m)} \quad , \quad t \geq 0, \\ \begin{pmatrix} R_t(t) \\ \tau_t(t) \end{pmatrix} &= \begin{pmatrix} R(t) S(t) \\ R(t) \lambda(t) \end{pmatrix} \quad , \quad t > 0, \\ \begin{pmatrix} R(0) \\ \tau(0) \end{pmatrix} &= \begin{pmatrix} I_d \\ 0 \end{pmatrix} \quad , \quad t = 0. \end{aligned}$$

The last two equations in (15) for the position $\gamma = (R, \tau)$ are decoupled from the other equations in (15). Therefore, the γ -equation can be solved in a postprocessing step. The γ -equation is called the [reconstruction equation](#) for the rotating wave. Since (v_\star, μ_\star) with $\mu_\star = (S_\star, \lambda_\star)$ satisfy

$$\begin{aligned} 0 &= A \Delta v_\star(\xi) + \langle S_\star \xi + \lambda_\star, \nabla v_\star(\xi) \rangle + f(v_\star(\xi)), \quad \xi \in \mathbb{R}^d, \\ 0 &= (v_\star - \hat{v}, (\xi_j D_i - \xi_i D_j) \hat{v})_{L^2(\mathbb{R}^d, \mathbb{R}^m)}, 0 \quad = (v_\star - \hat{v}, D_l \hat{v})_{L^2(\mathbb{R}^d, \mathbb{R}^m)}, \end{aligned}$$

we expect for stability reasons, that the solution (v, μ, γ) of (15) satisfies

$$(16) \quad v(t) \rightarrow v_\star, \quad \mu(t) = (S(t), \lambda(t)) \rightarrow \mu_\star = (S_\star, \lambda_\star) \quad \text{as } t \rightarrow \infty.$$

As an indicator for the convergence in (16) we check the quantities

$$(17) \quad \|v_t(\cdot, t)\|_{L^2(\mathbb{R}^d, \mathbb{R}^m)} \quad \text{and} \quad |\mu_t(t)|$$

at each time instance t during the computation. In fact, both of these quantities should be small ($\approx 10^{-16}$), since v_\star and μ_\star do not vary in time.

3.3 Numerical approximation of rotating waves via freezing method

Solving (1), (15) and (7) numerically, requires to truncate these equations to bounded domains. Let $\Omega \subset \mathbb{R}^d$ be a bounded open domain, then (1) must be satisfied for $x \in \Omega$, and equations (15) and (7) for $\xi \in \Omega$. To guarantee the well-posedness of these problems, we must equip the

equations with appropriate boundary conditions. Normally, we choose **homogeneous Neumann boundary conditions** (also known as **no-flux boundary conditions**), i.e.

$$\frac{\partial u}{\partial \mathbf{n}}(x) = 0, \quad x \in \partial\Omega, \quad \frac{\partial v}{\partial \mathbf{n}}(\xi) = 0, \quad \xi \in \partial\Omega.$$

In this context, $\partial\Omega$ denotes the **boundary** of Ω and $\overline{\Omega}$ the **closure** of Ω , e.g. $\Omega = B_R(x_0) = \{x \in \mathbb{R}^d \mid |x - x_0| < R\} \subset \mathbb{R}^d$ with $R > 0$ and $x_0 \in \mathbb{R}^d$ then $\partial\Omega = \partial B_R(x_0) = \{x \in \mathbb{R}^d \mid |x - x_0| = R\}$ and $\overline{\Omega} = \overline{B_R(x_0)} = \{x \in \mathbb{R}^d \mid |x - x_0| \leq R\}$. Numerically, we solve the following equations:

Step 1: (Nonfrozen Equation)

$$(18) \quad \begin{aligned} u_t(x, t) &= A\Delta u(x, t) + f(u(x, t)) \quad , \quad x \in \Omega, \quad t \in (0, T_1], \quad d \geq 2, \\ u(x, 0) &= u_0(x) \quad , \quad x \in \overline{\Omega}, \quad t = 0, \\ \frac{\partial u}{\partial \mathbf{n}}(x, t) &= 0 \quad , \quad x \in \partial\Omega, \quad t \in [0, T_1]. \end{aligned}$$

First, we determine the solution u of (18). The quantities A , f , u_0 , Ω and T_1 are given.

Step 2: (Frozen Equation)

$$(19) \quad \begin{aligned} v_t(\xi, t) &= A\Delta v(\xi, t) + \langle S(t)\xi + \lambda(t), \nabla v(\xi, t) \rangle + f(v(\xi, t)) \quad , \quad \xi \in \Omega, \quad t \in (0, T_2], \\ v(\xi, 0) &= v_0(\xi) \quad , \quad \xi \in \overline{\Omega}, \quad t = 0, \\ \frac{\partial v}{\partial \mathbf{n}}(x, t) &= 0 \quad , \quad \xi \in \partial\Omega, \quad t \in [0, T_2], \\ 0 &= (v(\cdot, t) - \hat{v}, (\xi_j D_i - \xi_i D_j)\hat{v})_{L^2(\Omega, \mathbb{R}^m)} \quad , \quad t \in [0, T_2], \\ 0 &= (v(\cdot, t) - \hat{v}, D_l \hat{v})_{L^2(\Omega, \mathbb{R}^m)} \quad , \quad t \in [0, T_2], \\ \begin{pmatrix} R_t(t) \\ \tau_t(t) \end{pmatrix} &= \begin{pmatrix} R(t)S(t) \\ R(t)\lambda(t) \end{pmatrix} \quad , \quad t \in (0, T_2], \\ \begin{pmatrix} R(0) \\ \tau(0) \end{pmatrix} &= \begin{pmatrix} I_d \\ 0 \end{pmatrix} \quad , \quad t = 0. \end{aligned}$$

Then, we determine the solution (v, μ, γ) with $\mu = (S, \lambda)$ and $\gamma = (R, \tau)$ of (19). The quantities A , f , v_0 , \hat{v} , Ω and T_2 are given. The final time T_2 may be different to the end time T_1 from (18). The template function is often chosen as $\hat{v}(\xi) = u_0(\xi)$ or $\hat{v}(\xi) = u(\xi, T_1)$, where $u(\cdot, T_1)$ denotes the solution of (18) at the end time T_1 . Sometimes one must solve (18) to obtain a suitable template function \hat{v} . Also the initial data v_0 is often chosen as $v_0(\xi) = u_0(\xi)$ or $v_0(\xi) = u(\xi, T_1)$. The end time T_2 in (19) is often chosen such that the values of the quantities $\|v(\cdot, t)\|_{L^2(\Omega, \mathbb{R}^m)}$ and $|\mu_t(t)|$, cf. (17), are near 10^{-16} .

Step 3: (Eigenvalue Problem)

$$(20) \quad \begin{aligned} \lambda w(\xi) &= A\Delta w(\xi) + \langle S_\star \xi + \lambda_\star, \nabla w(\xi) \rangle + Df(v_\star(\xi))w(\xi) \quad , \quad \xi \in \Omega, \\ \frac{\partial w}{\partial \mathbf{n}}(\xi) &= 0 \quad , \quad \xi \in \partial\Omega. \end{aligned}$$

Finally, we determine (a prescribed number neig of) eigenvalues λ and associated eigenfunctions w of (20). The quantities A , $\mu_\star = (S_\star, \lambda_\star)$, v_\star , f , Ω and neig are given. The profile v_\star and the velocity $\mu_\star = (S_\star, \lambda_\star)$ come actually from a simulation, more precisely we set $\mu_\star := \mu(T_2)$ and $v_\star(\xi) := v(\xi, T_2)$, where $\mu(T_2)$ and $v(\cdot, T_2)$ denote two components of the solution of (19) at the end time T_2 .